Beam Background at Super-KEKB/Belle

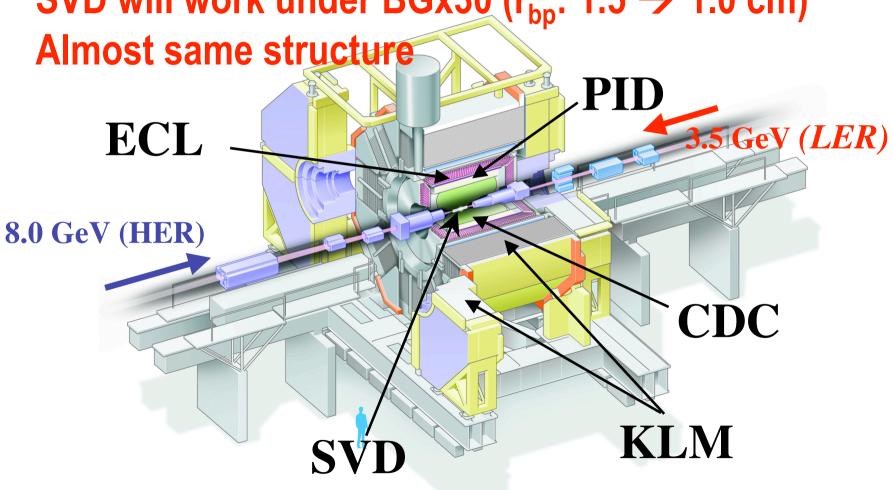
Osamu Tajima (KEK, IPNS) for Belle collaboration Feb 22, 2005 KEKB review

Outline

- Current Status
- Expectation & Reduction
- Summary

Belle Detector

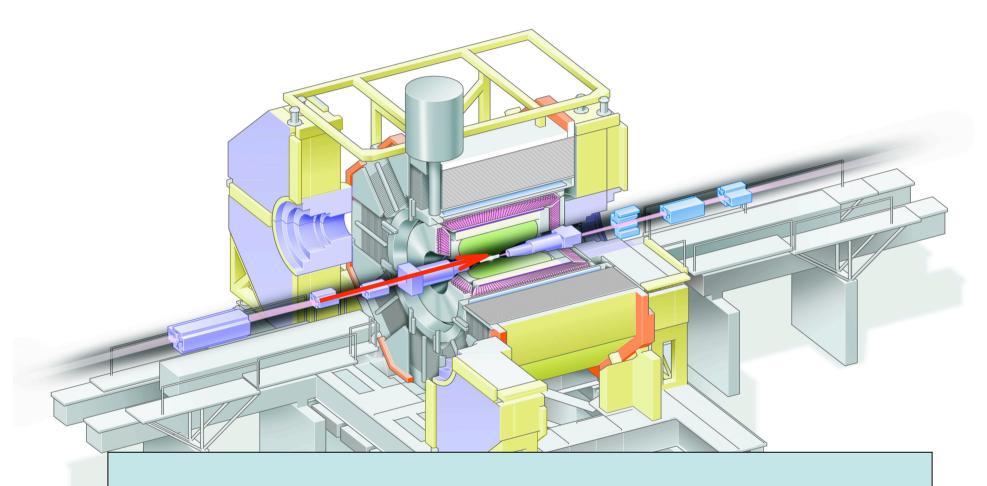
Detector will be upgrade to work under BGx20 SVD will work under BGx30 (r_{bp} : 1.5 \rightarrow 1.0 cm)



Beam backgrounds should be concern

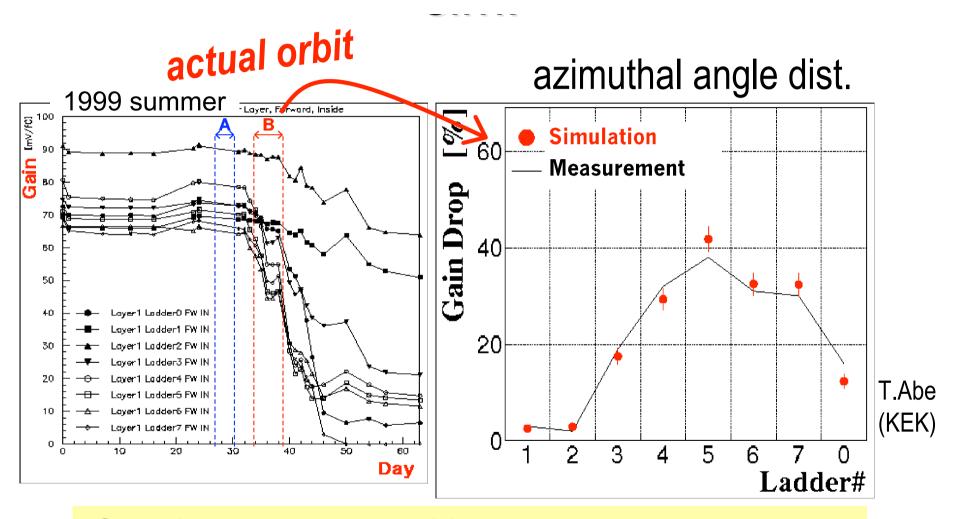
- Synchrotron Radiation (SR) photons
 - > generated in upstream magnets
 - > generated in downstream (QCS) magnet
- Shower caused by spent particles
 - beam-gas scattering
 - Touschek scattering
- Radiative Bhabha origin
 - > Neutrons from downstream beamline
 - > Showers caused by over bend beams

SR from upstream magnets



SVD1.0 killer source (1999)

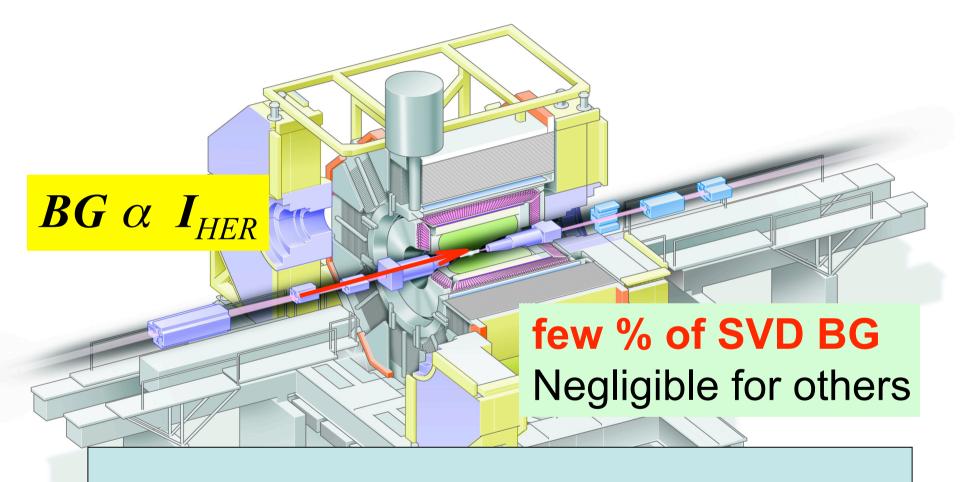
Past troub'



Steering angle was critical

No more trouble w/ limitation of steering magnet

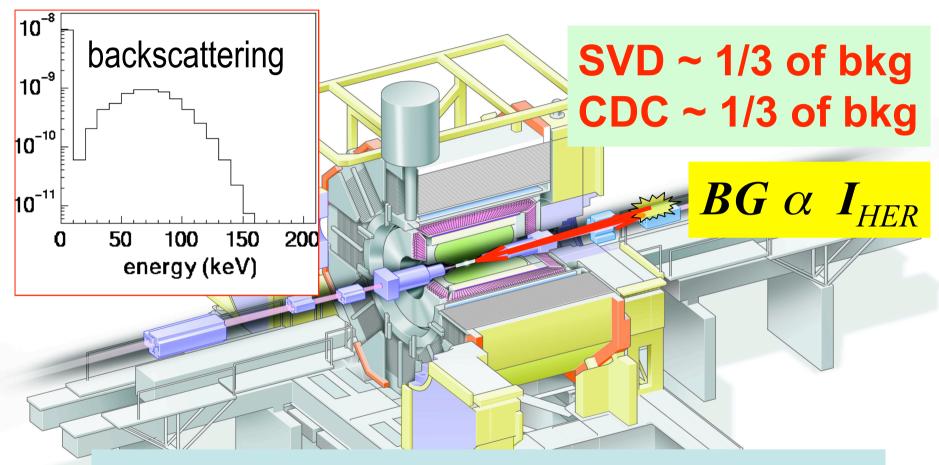
SR from upstream magnets



SVD1.0 killer source (1999)

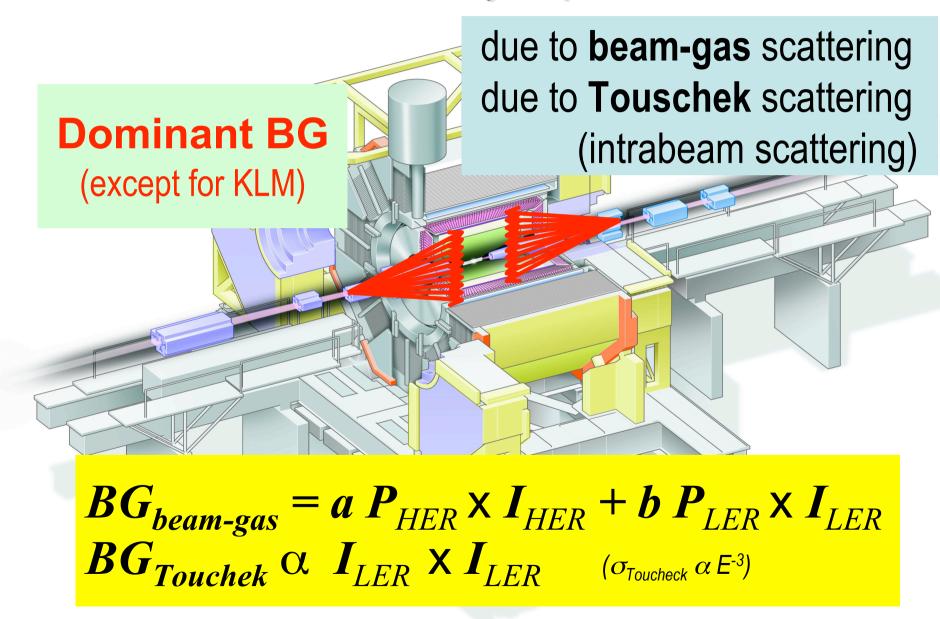
→ Not serious after limitation of steering

SR, downstream magnet (QCS) origin

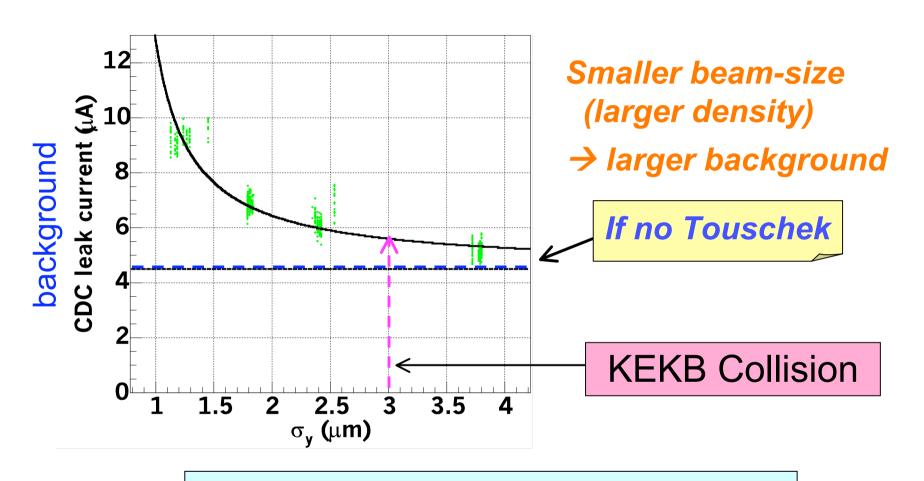


- Downstream final focus magnet (QCS) generate high energy SR (E_{crit} ~ 40 keV)
- 2. SR photons are scattered at downstream chamber (~9m)
- 3. Backscattering photons enter to the detector ($E_{eff} \sim 100 \text{ keV}$)

Shower caused by Spent Particles

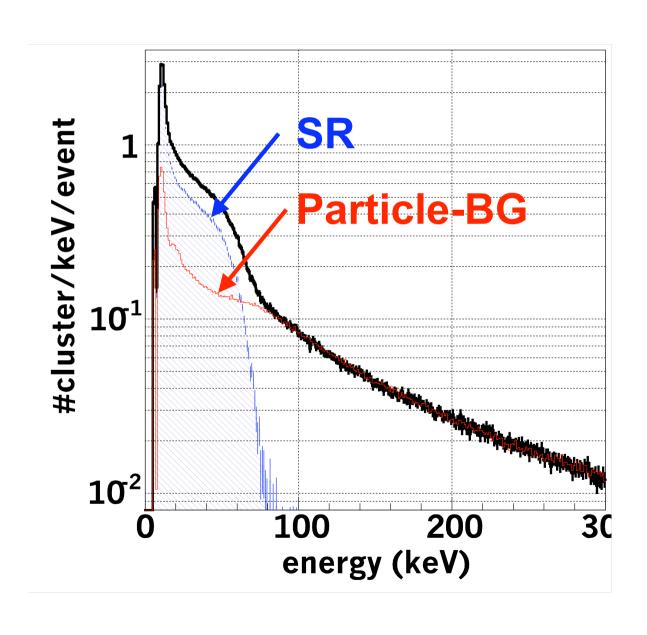


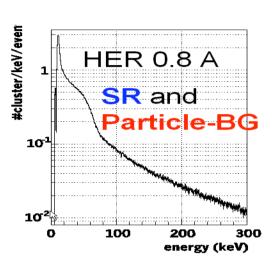
Study for Contribution of Touschek

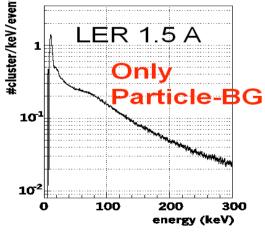


Contribution ~ 20 % of LER BG from particle showers

SVD BG Extraction

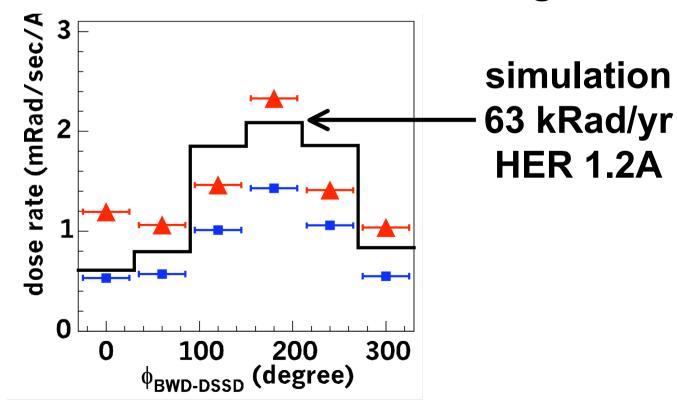




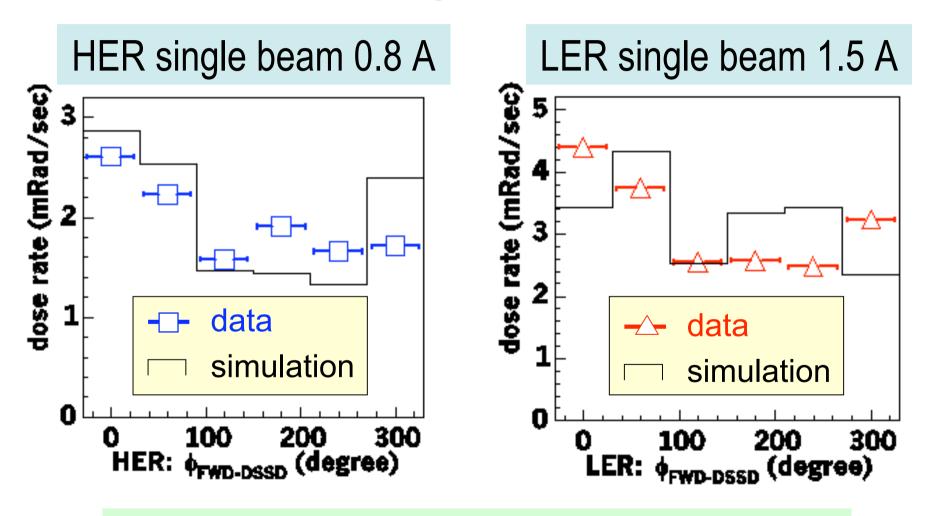


Azimuthal angle dist. of SR at 1st layer of SVD

- ▲ Single-Bunch 15 mA (trigger-timing is adjusted)
- Total 0.8 A w/ 1284 bunch (random timing)
- ☐ Simulation for **SR-backscattering**

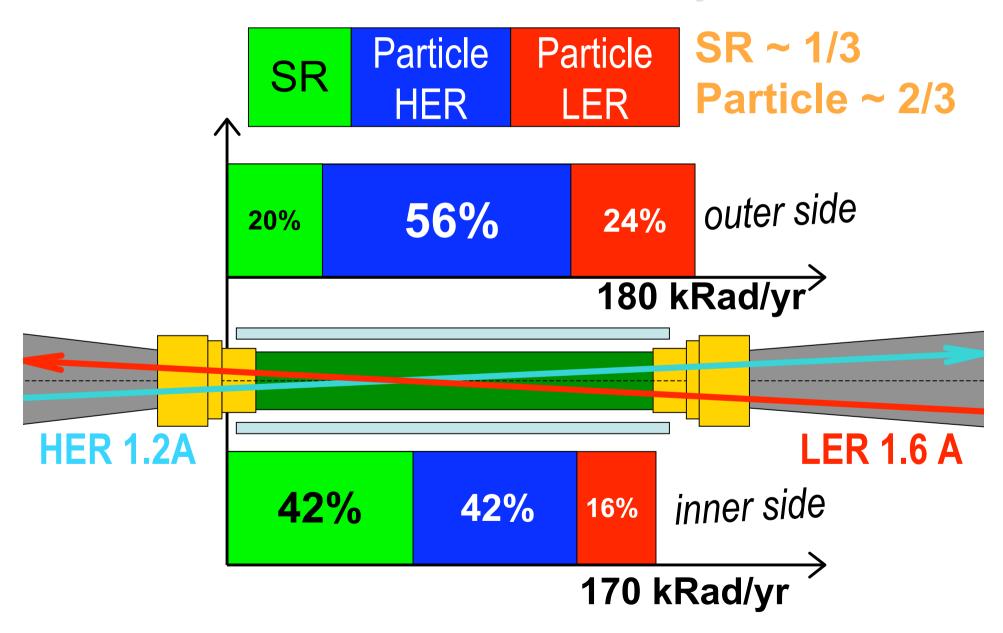


Azimu. angle dist. of Shower particles at 1st layer of SVD

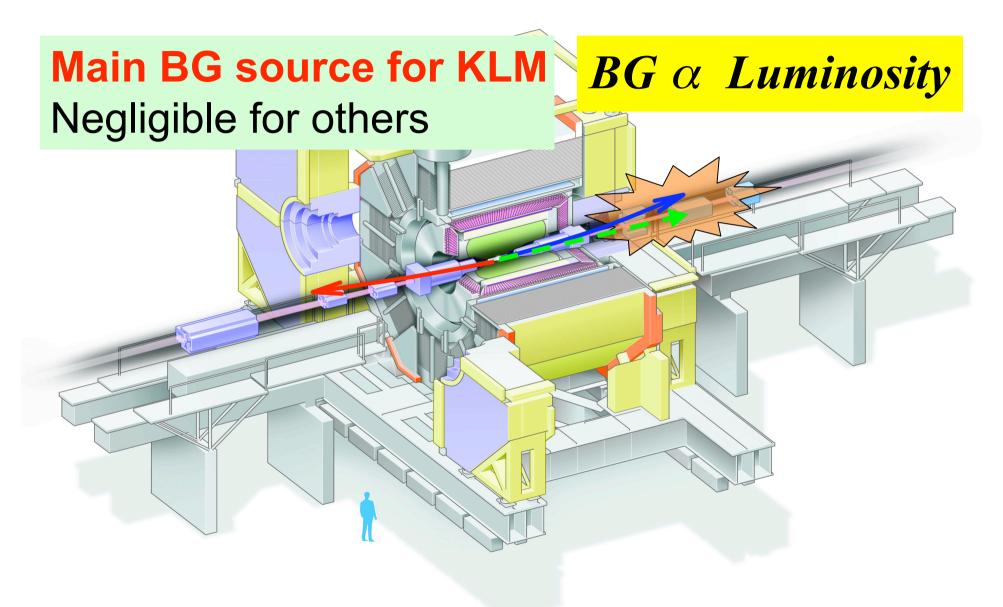


Almost consistent with simulation ~40% accuracy

BG of SVD at 1st layer

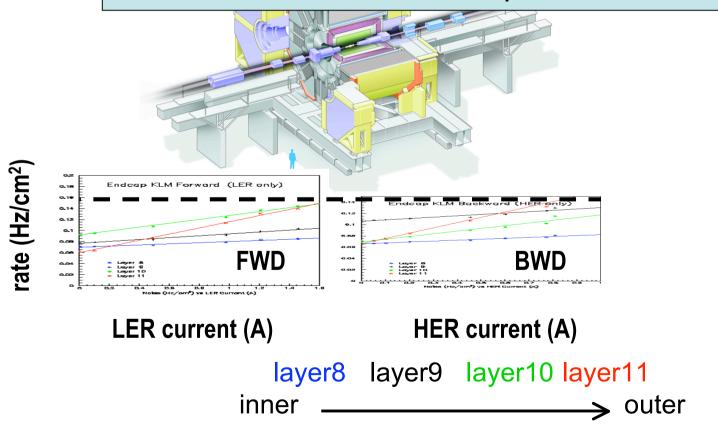


Radiative Bhabha origin



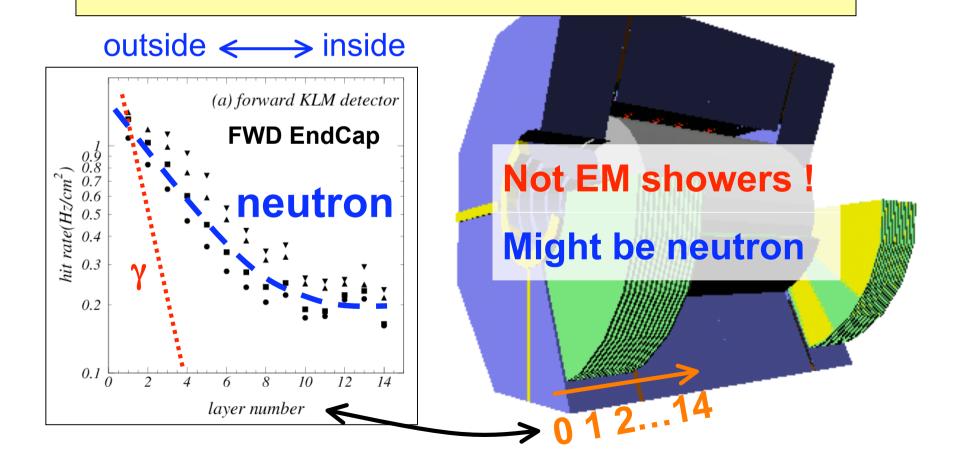
KLM: EndCap

Main bkg source is luminosity origin Luminosity component ~ 75 % Beam current component ~ 25 %



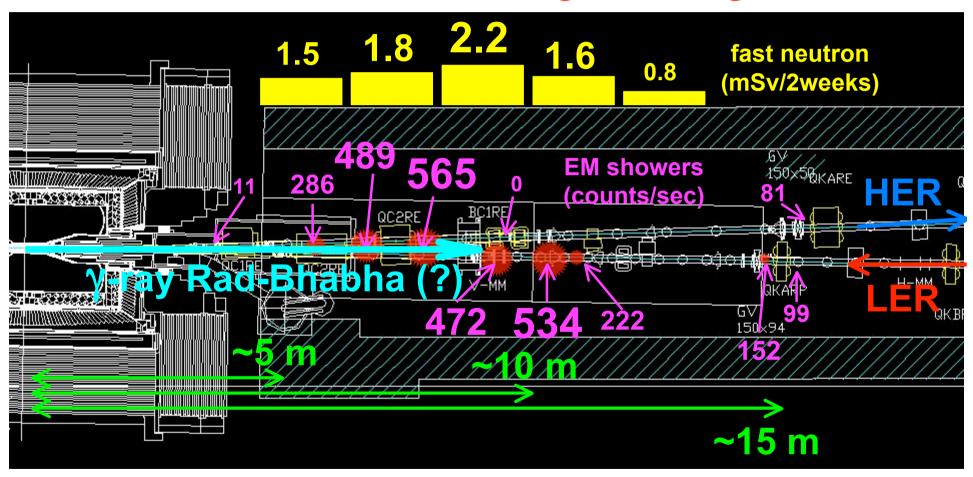
What is KLM bkg?

Strong correlation with Luminosity



Where is neutron source?

Radiative-Bhabha might be origin

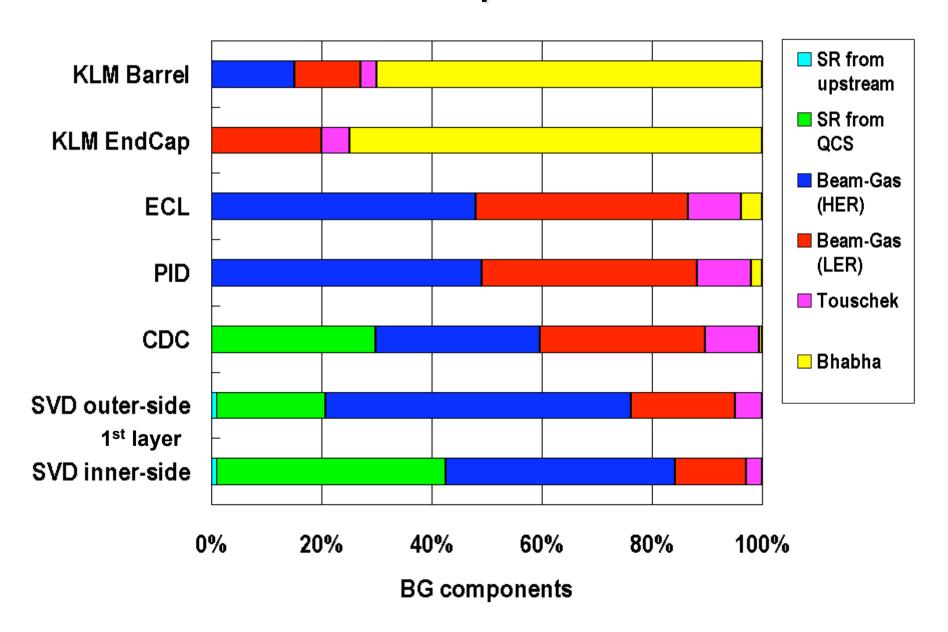


Apr, 2003 (by Tawara, Nakamura)

Oct 27,2004 (by K.Abe & T.Sarangi)

LER / HER = 1630 / 1160 mA

BG components



BG Estimation & Reduction for Super-KEKB/Belle

Machine parameters for BG estimation

	KEKB		SuperKEKB	
	LER	HER	LER	HER
I [A]	1.6	1.2	9.4	4.1
$\beta_{x}^{*}[cm]$	33	33	20	20
$\beta_{y}^{*}[mm]$	10	10	3	3
L [/cm ² /sec]	1.3×10^{34}		2.5×10^{35}	
σ_{y} [um]	3	3	1.5	1.5
σ_{l} [mm]	5	5	3	3
Ave. Vacuum (10-7 Pa)	1.25	1.25	5 (2.5)	5 (2.5)
φ [mrad]	11		15	
SR power : QCSR (kW)	29		179	
Run time (days/year)	200		200	

Backscattering of QCS-SR

- Larger current and critical energy gives larger QCS-SR wattage → x6 (179 kW)
- Assume that scattered position will be same as current KEKB (~9m)
 - > IR chamber design is not finalized
- 6 times higher than current SR-BG
 Dose at 1st layer of SVD
 - > 400 kRad/yr at inner side of the ring
 - > 130 kRad/yr at outer side of the ring

Shower of Spent Particles

Beam-Gas scattering origin

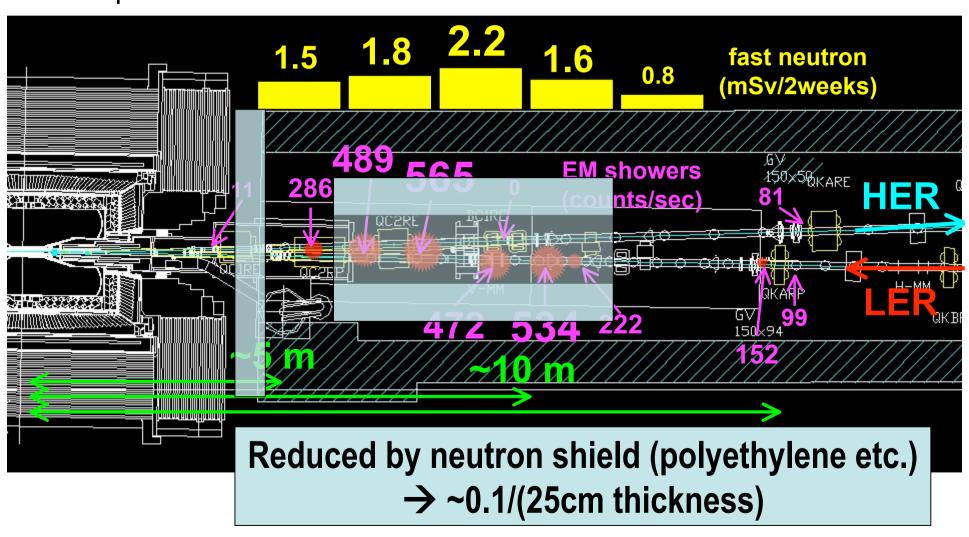
- \rightarrow HER / LER currents: 1.2/1.6 A \rightarrow 4.1/9.4 A
- > x4 Worse Vacuum : $1 \sim 1.5 \times 10^{-7} \text{ Pa} \rightarrow 5 \times 10^{-7} \text{ Pa}$
- > x13.6 (HER), x23.6 (LER) higher contribution

Touschek origin

- > smaller beam size ~ 1.5um/3um
- > shorter bunch length ~ 3mm/5mm
- > higher bunch current ~ 1.5 times
- > many #bunch ~ 3.9 times
- > currently just 10% of total BG
- > x97 times higher contribution

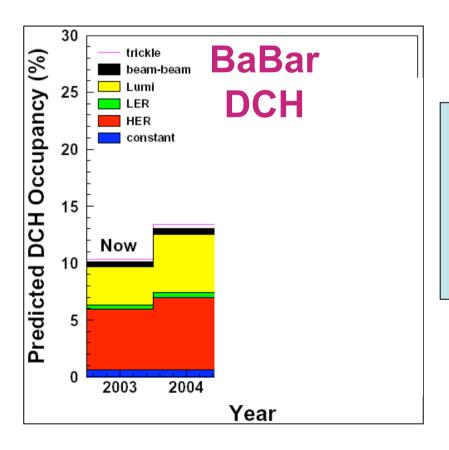
Rad. Bhabha: KLM

 $L_{\text{super-KEKB}}$ / L_{KEKB} = 25 / 1.3 ~ 19 times higher

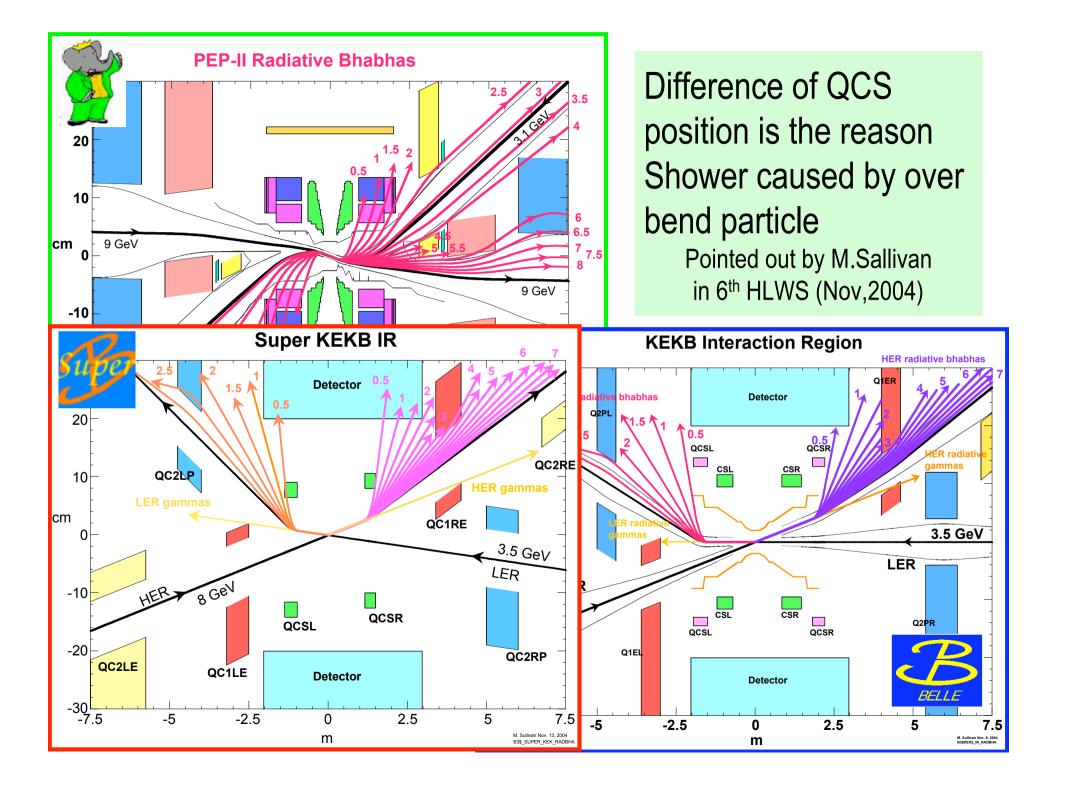


Radiative Bhabha: inner detectors

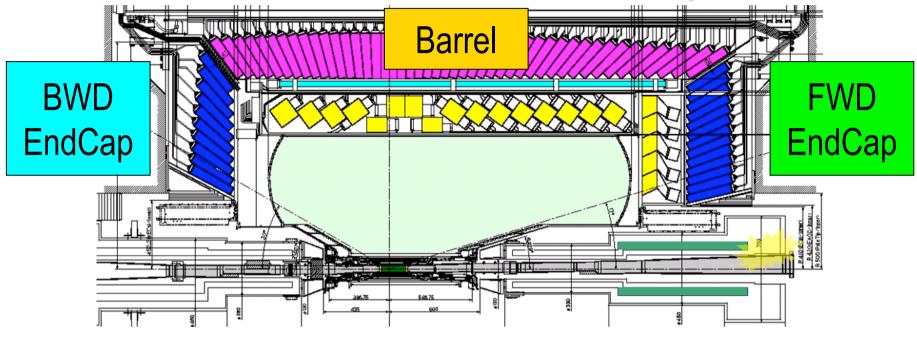
 Actually, BaBar has large BG for inner detectors while it is negligible at Belle

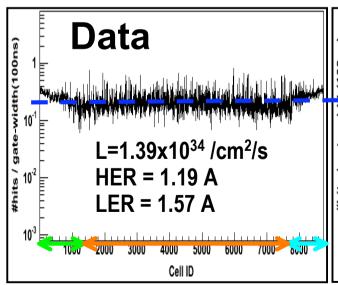


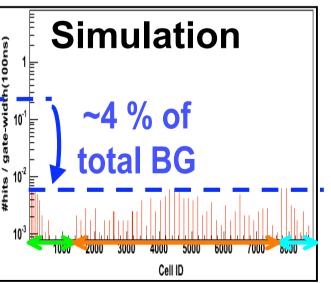
We should consider because higher lum gives higher BG



Rad. Bhabha BG at ECL, NOW

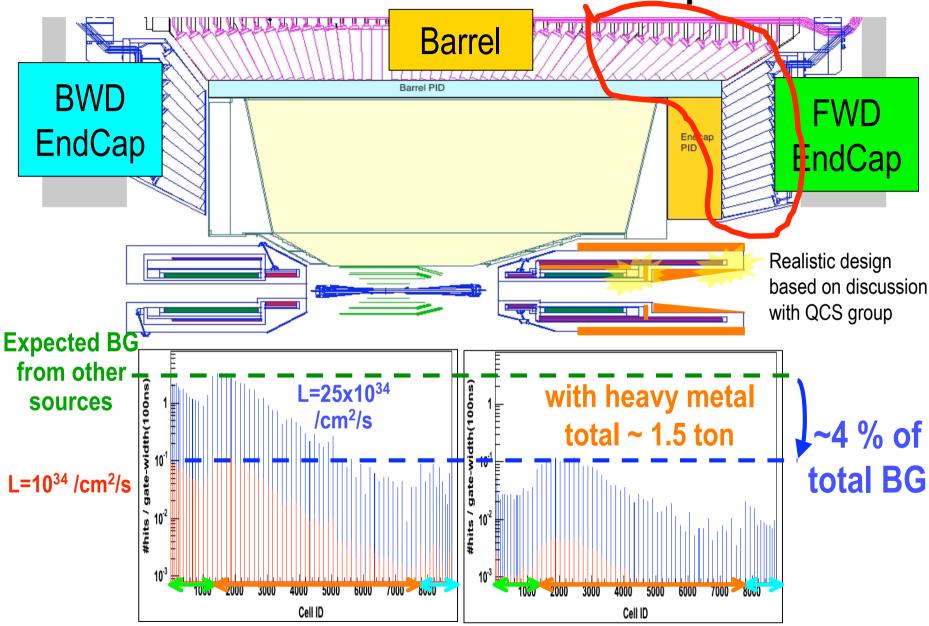


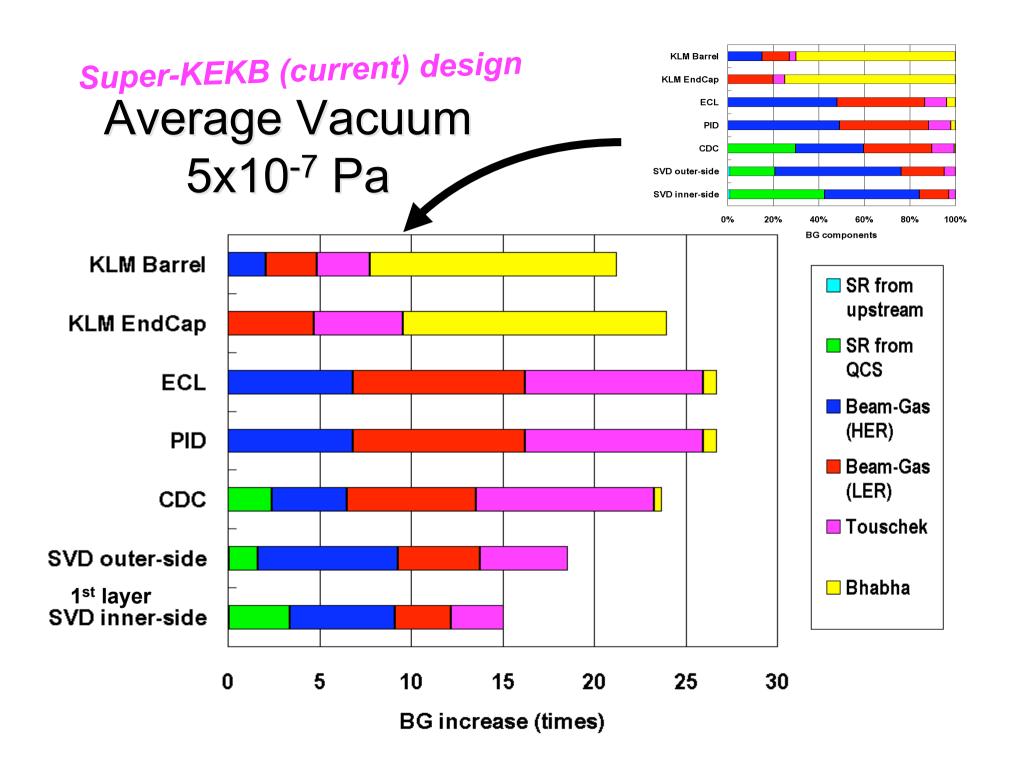


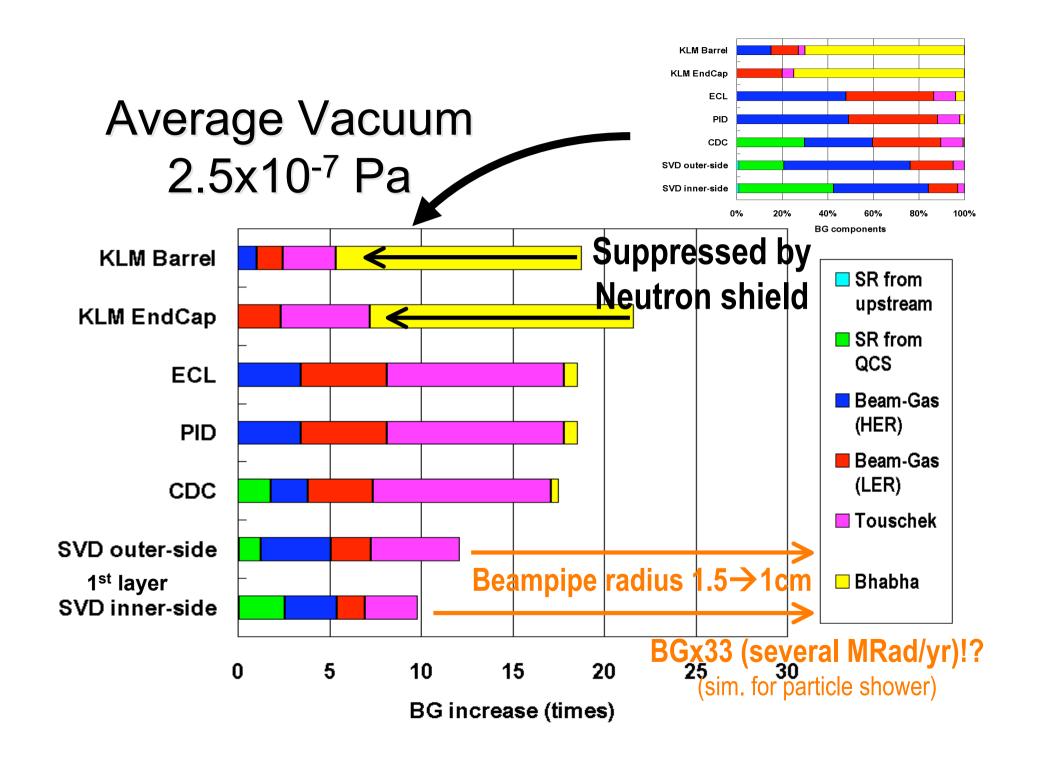


Asuuming L=10³⁴ /cm²/s

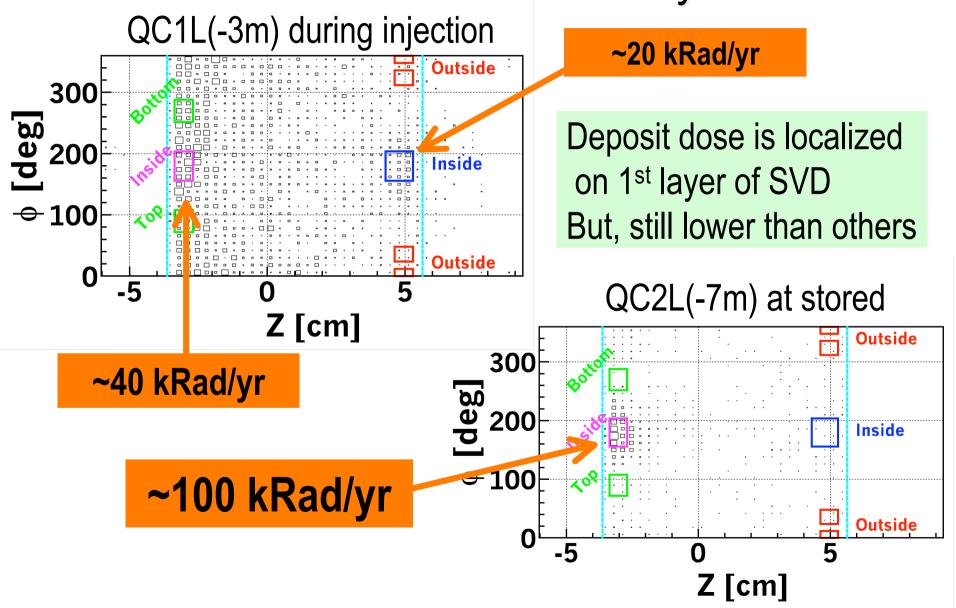
Rad. Bhabha BG sim. for Super-KEKB







SR from upstream magnets simulation for SVD 1st layer



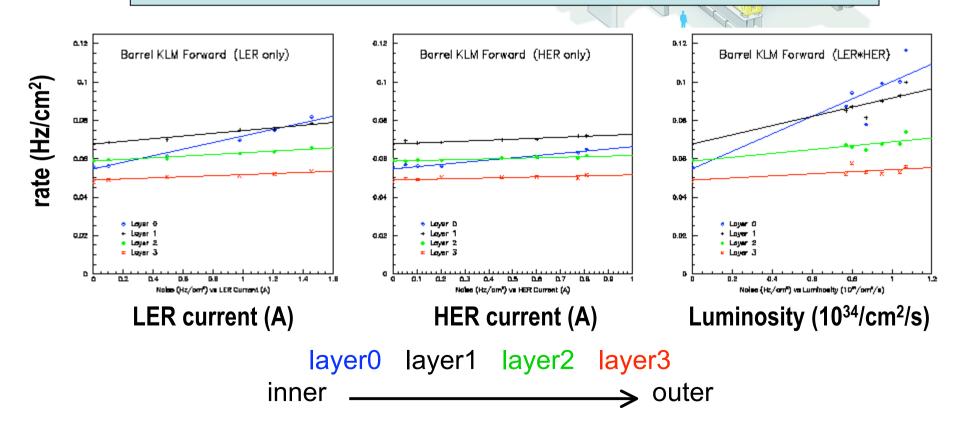
Summary

- Backscattering of QCS-SR is not serious, but strongly depends on IR chamber configuration
- Vacuum level is very important
 - > Original design (5x10⁻⁷ Pa) is serious \rightarrow BGx25
 - > w/ further effort $(2.5x10^{-7} Pa) \rightarrow BGx18 \leftarrow -30\%$
- Increasing of Touschek origin BG
 - > Smaller bunch size & higher bunch currents are reason
 - Might be reduced by further study
- Radiative Bhabha origin BG can be suppressed
- Beampipe radius 1.5cm → 1cm
 - > Further simulation study of shower particles into SVD is important

backup

KLM: Barrel

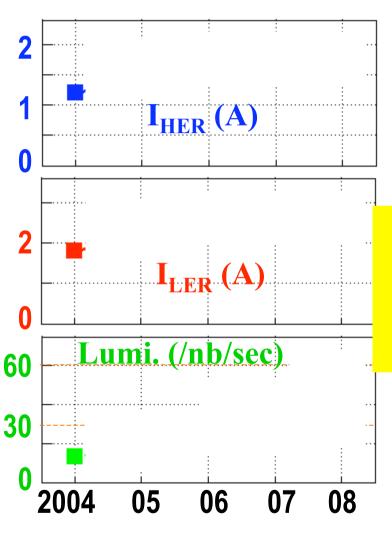
Main bkg source might be luminosity origin Luminosity component ~ 70% Beam current component ~ 30 %



Now

- HER / LER = 1.2 / 1.8 A
- Luminosity ~ 13 /nb/sec (1.3x10³⁴ /cm²/sec)
- Which is dominant background?
 - SVD : SR ~ 1/3, Spent particles ~ 2/3
 - CDC : SR ~ 1/4, Spent particles ~ 3/4
 - $(HER \sim 1/4, LER \sim 2/4)$
 - KLM EndCap : L ~ 75 %, Spent particles ~ 25 %
 - Barrel : $L \sim 70 \%$, Spent particles $\sim 30 \%$
 - Others : Spent particles might be dominat

Near Future Prospects of KEKB



Now (2004)

L = 13 / nb/sec

HER / LER = 1.2 / 1.8 A

RF limits beam-current

RF repair & minor upgrade (2005)

within 3~4 years ...

Beam current: x 1.7 higher

Luminosity: x 4.6 higher

Crab cavity (2006) HER&LER

→ twice luminosity !?

Minor upgrades (2007-2008)

→ L = 60/nb/sec HER / LER = 2 / 3 A

Near Future

```
SR (\alpha I<sub>HER</sub>)
Spent particles (\alpha I<sup>2</sup>)
Luminosity origin (\alpha L)
```

```
SVD outer x 2.6 inner x 2.3
```

```
CDC x 2.5
```

KLM EndCap x 4.2

Barrel x 4.1

Others x 2.8

```
If we can keep good vacuum
( optimistic case )
most of detectors x 1.7
KLM x 3.9

lower limit !?
```

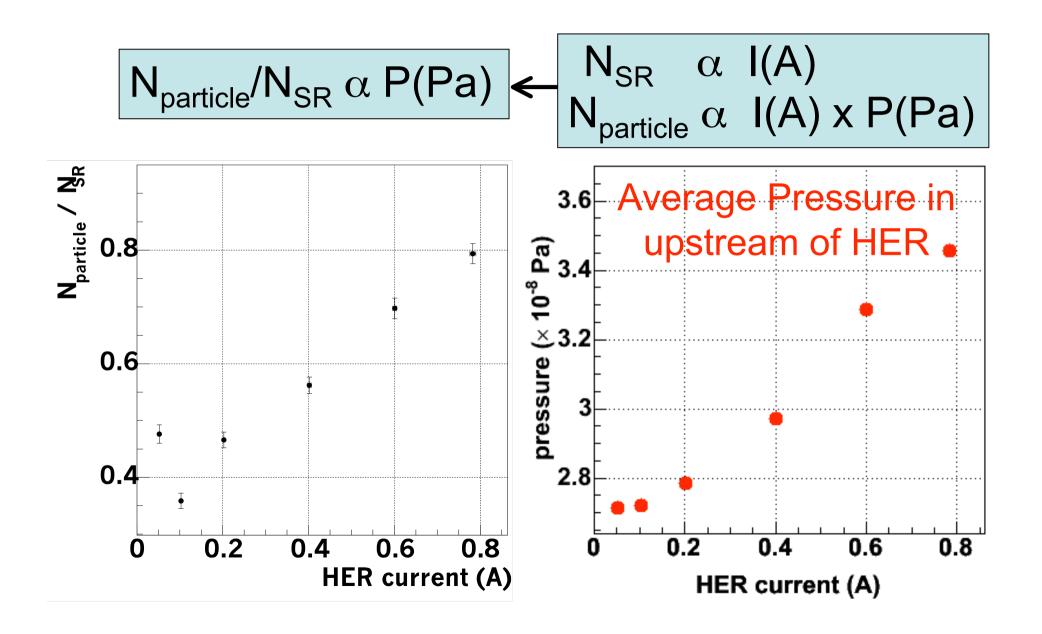
3 times bkg is good target value for near future

Summary

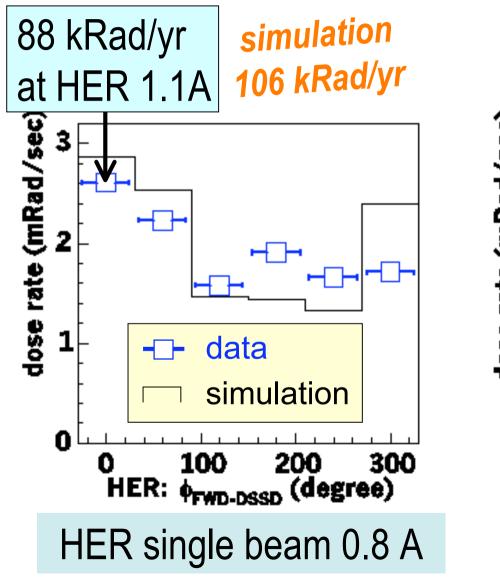
	2004	\rightarrow	2008
Lumi.	13 /nb/s	\rightarrow	60 /nb/s
HER	1.2 A	\rightarrow	2 A
LER	1.8 A	\rightarrow	3 A
SVD	Bkg	\rightarrow	x 1.7 ~ 2.6
CDC	Bkg	\rightarrow	x 1.7 ~ 2.5
KLM	Bkg	\rightarrow	x 3.9 ~ 4.2
Others	Bkg	\rightarrow	x 1.7 ~ 2.8

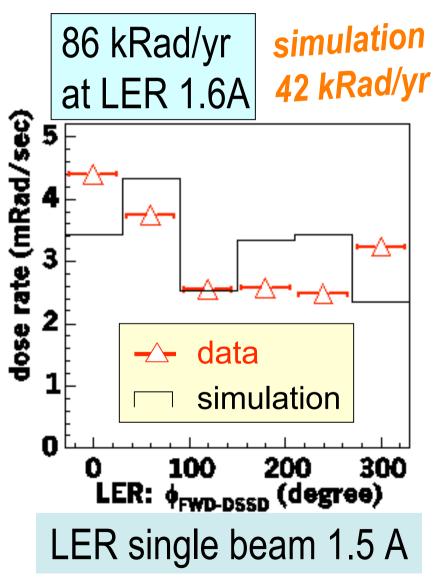
Higher bkg study must be needed for ~3 year later KLM bkg reduction must be start as soon as possible

Beam Current dep. of Vacuum



Azimu. angle dist. of Shower particles

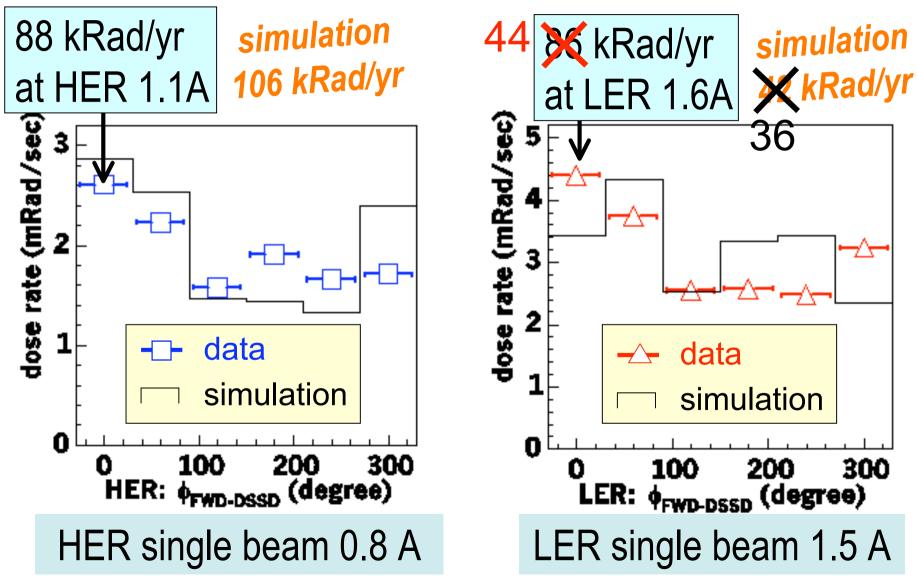




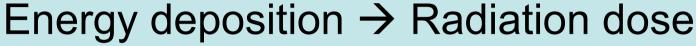
Azimu. angle dist. of Shower particles

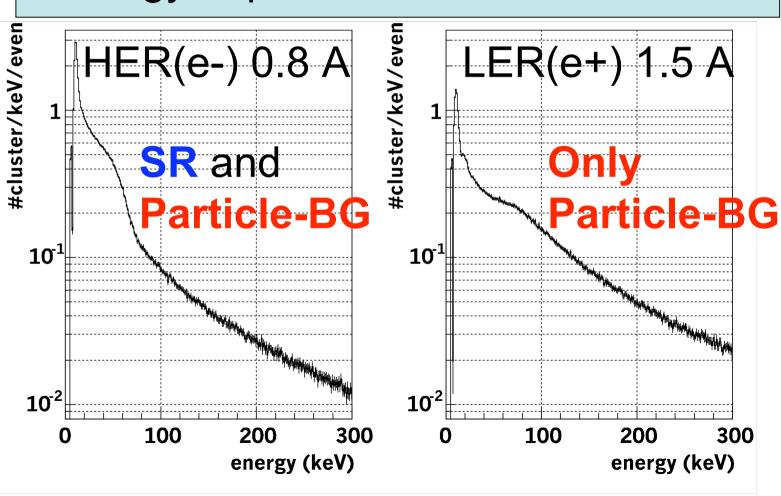
w/ correction of Touschek

300

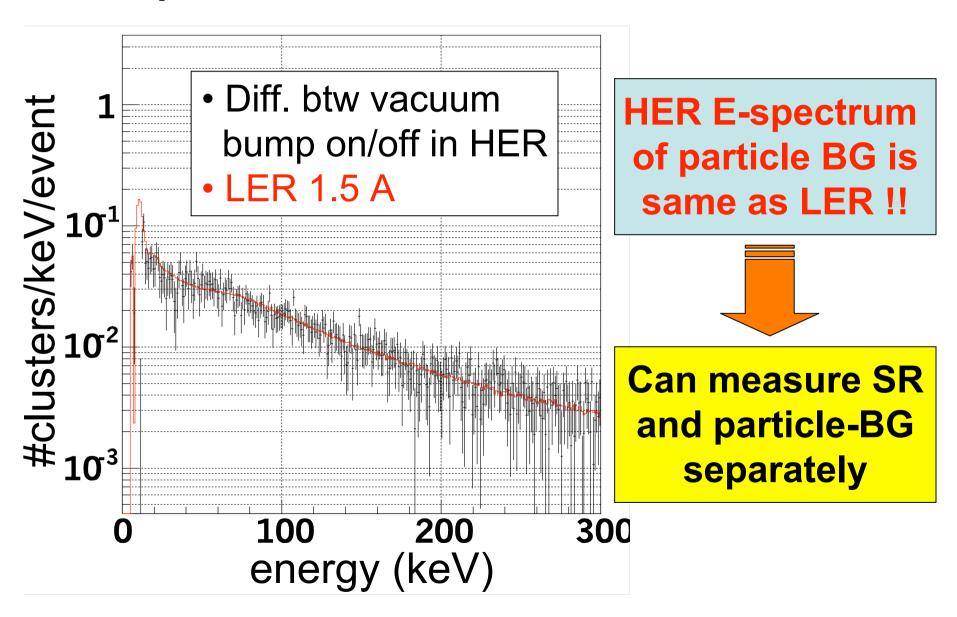


SVD Cluster Energy Spectra in Single Beam Run (1st layer)





E-spectrum of HER Particle-BG



Variation of Pressures at Pump

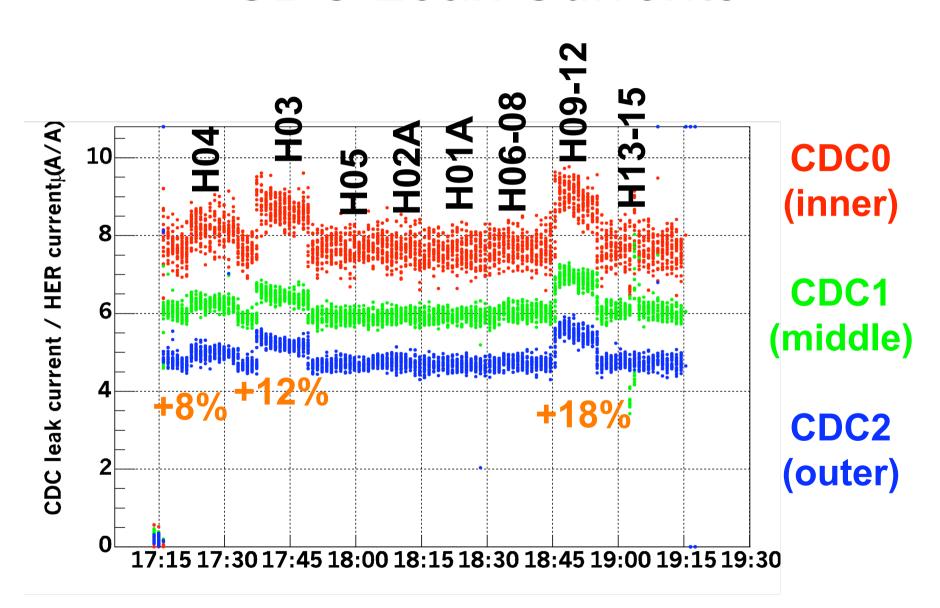
pressure (x10-8 Pa)

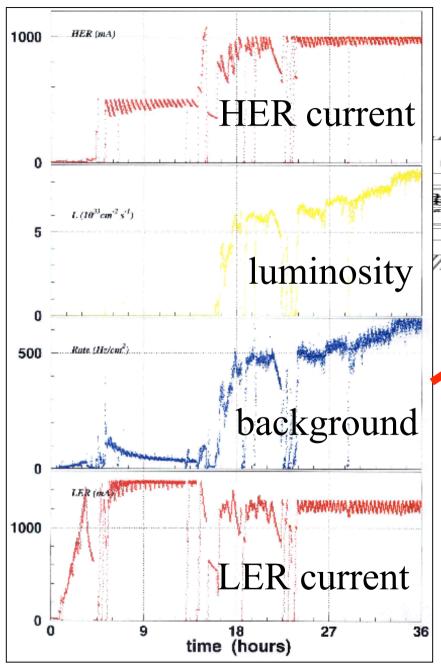
H01A (3.5m)	1.9 → 3.3 (x1.7)
H02A (7m)	6.7 → 14 (x2.1)
H03 (15m)	15 → 35 (x2.4)
H04 (25m)	4.8 → 13 (x2.7)
H05 (35m)	1* → 6.8 (x6.8)
H06-08 (~80m)	1* → 13 (x13)
H09-12 (~120m, ARC)	10 → 259 (x2.6)
H13-15 (~160m)	4.5 → 107 (x24)

monitor limit *

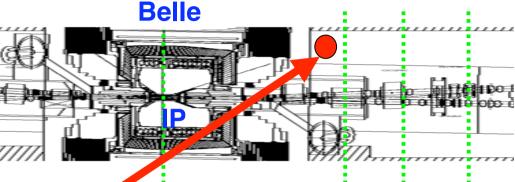
Maximum value

CDC Leak Currents





鉛ブロック(5cm厚)で囲まれ たプラスチックシンチレー ター検出器



ビーム電流だけでなく、ルミノシティーに比例したバックグランド(中性子と思われる)が大量に存在する

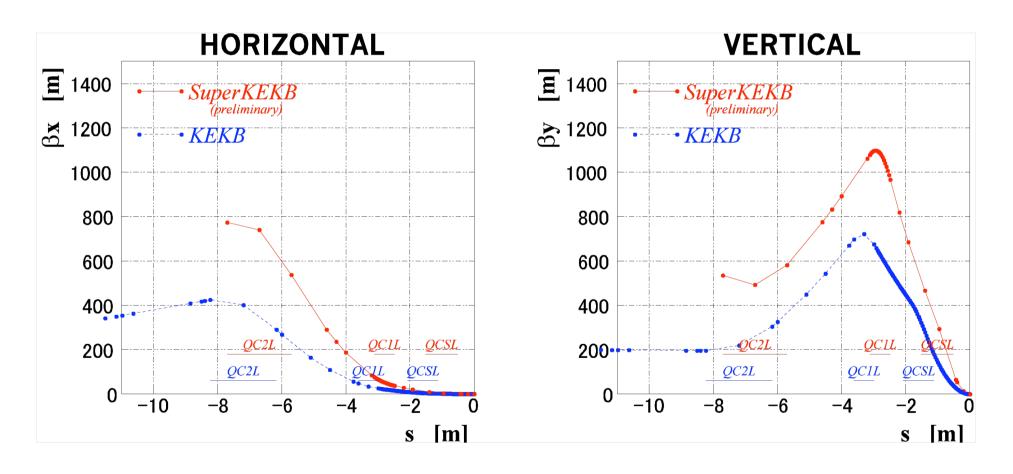
SuperKEKBではルミノシ ティーは50倍になる!!

Super-KEKB Machine Parameters

	KEKB	SuperKEKB
Horizontal emittance	24 nm	33 nm
x-y coupling	~ 3 %	6.4 %
$\beta_{x}^{*} / \beta_{y}^{*}$ (cm)	63 / 0.7	30 / 0.3
Beam current	1.1 A	4.1 A
Crossing angle	22 mrad	30 mrad
HER-beampipe tilt	11 mrad	15 mrad

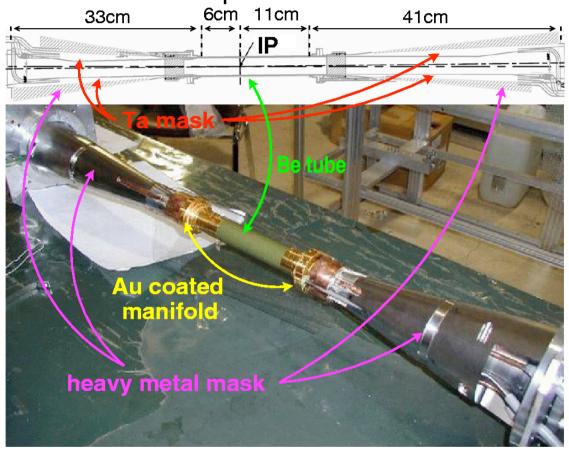
HER β Functions

Based on IR_HER6.sad

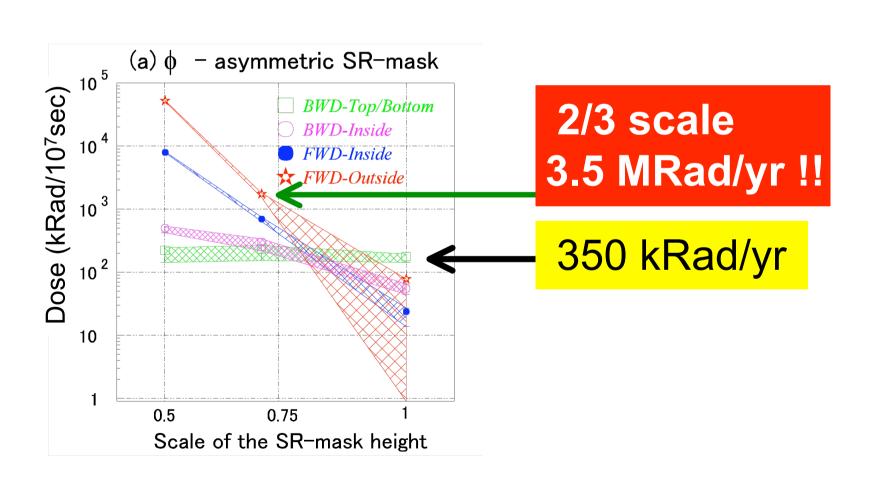


Beampipe in Super-Belle

• Scaled to 2/3 (r_{bp} 1.5 \rightarrow 1.0 cm)



→ Can we also scale SR-mask height?



Mask height should be same as r_{bp} =1.5 cm case (~2.5 mm height)

